

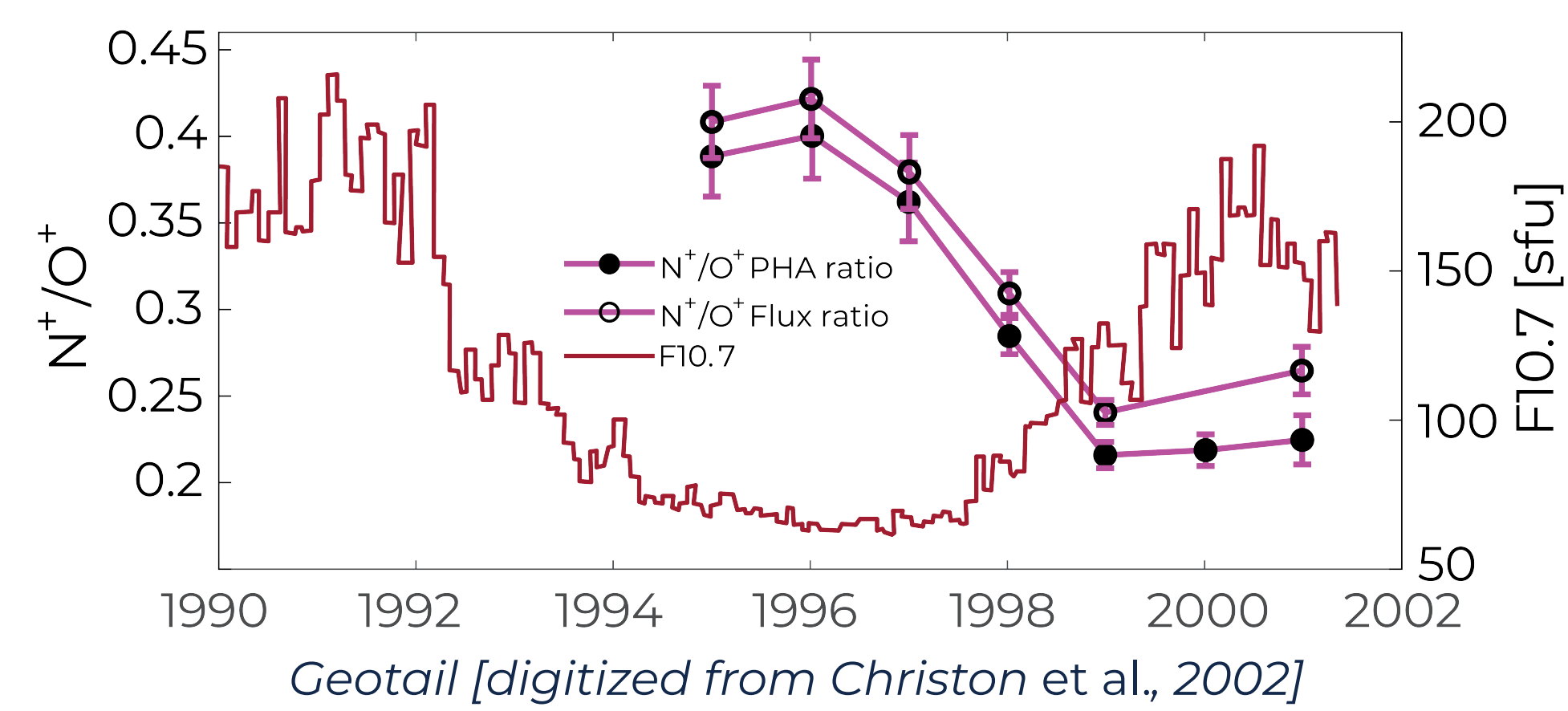
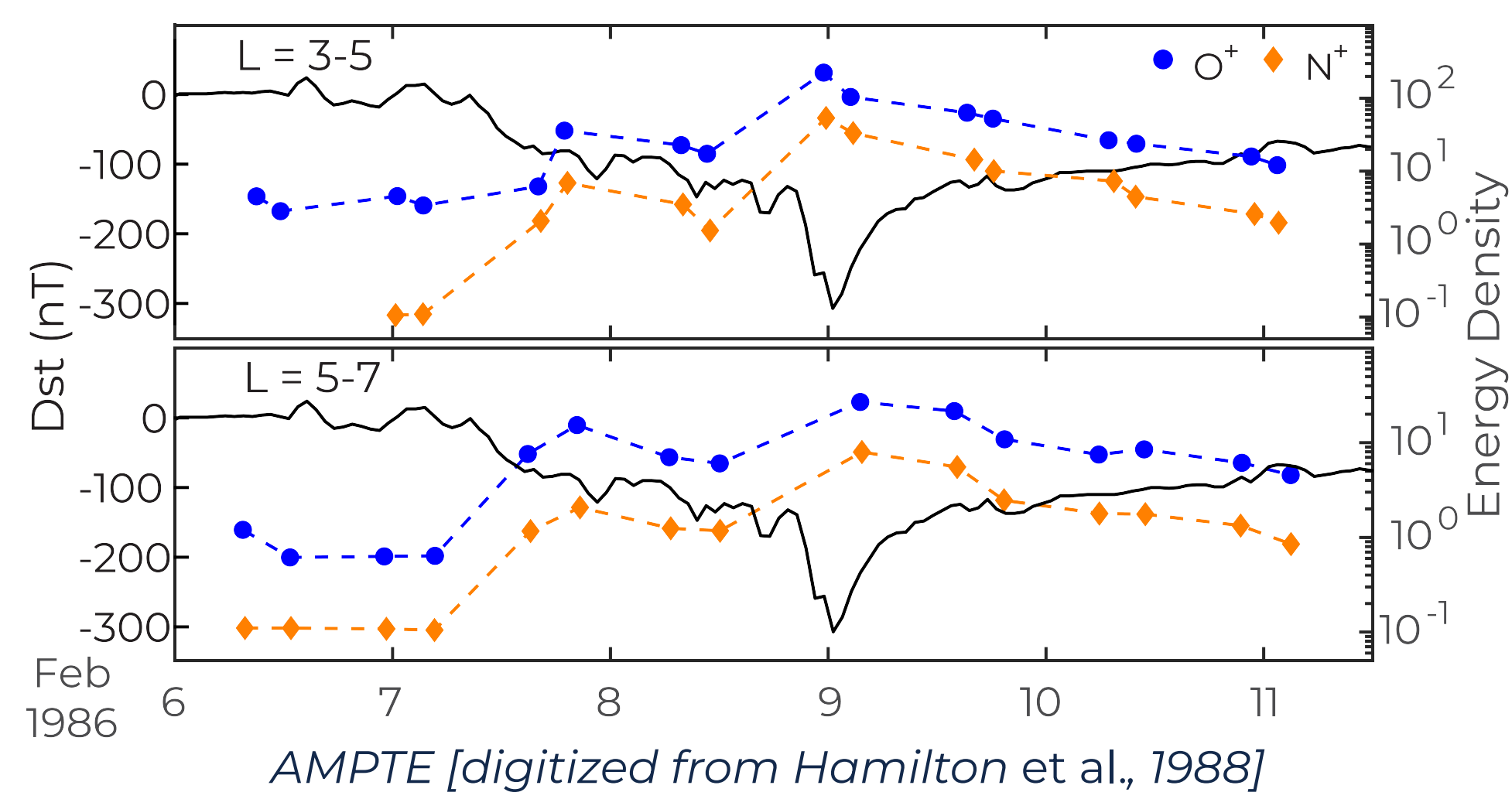
Impact of Plasma Mass Density on the Magnetosphere Configuration

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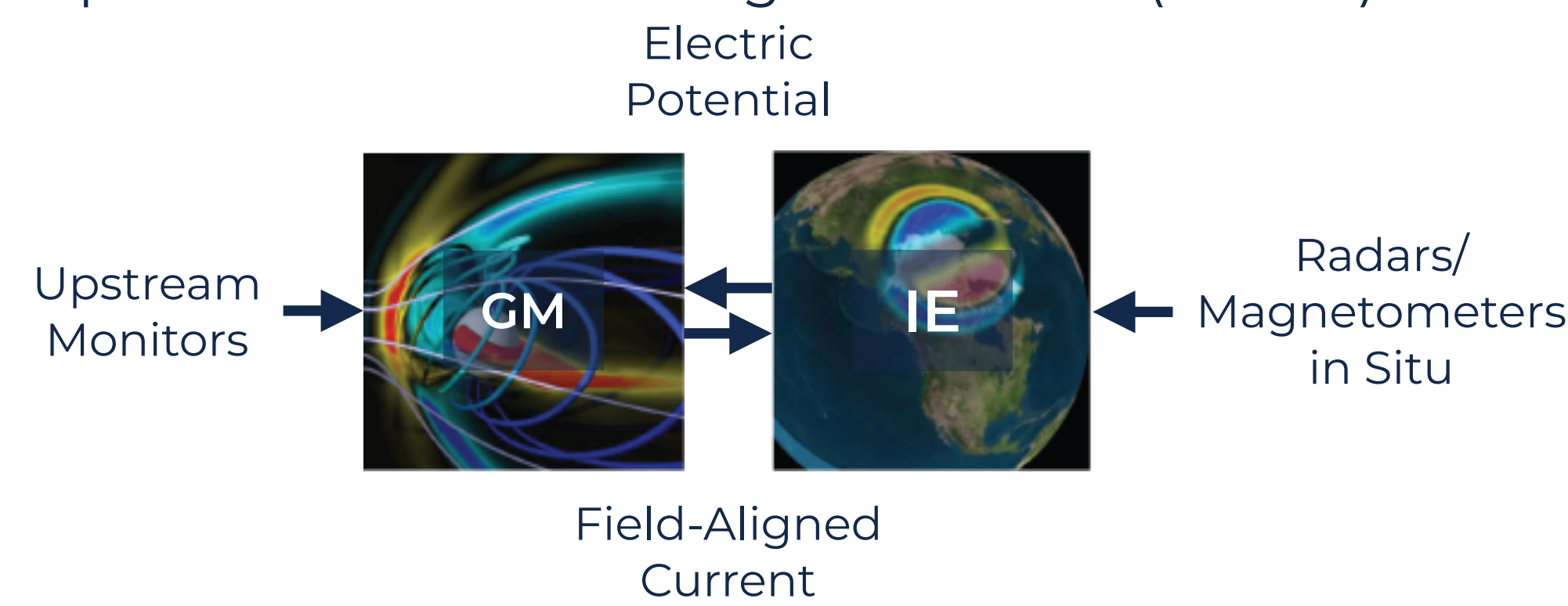
MOTIVATION

- Heavy ions regulate many magnetosphere processes. [Lin *et al.*, 2020]
 - N^+ ions have been observed in the ionosphere & magnetosphere since the 1960s, but instrumentation limitations have hindered further understanding of their dynamics. [Chappell *et al.*, 1982; Lin *et al.*, 2022]
 - $n(N^+)$ can be comparable to $n(O^+)$ in the inner magnetosphere. [Craven, 1993]
- What is the impact of inner magnetospheric plasma mass density on the global magnetosphere configuration?



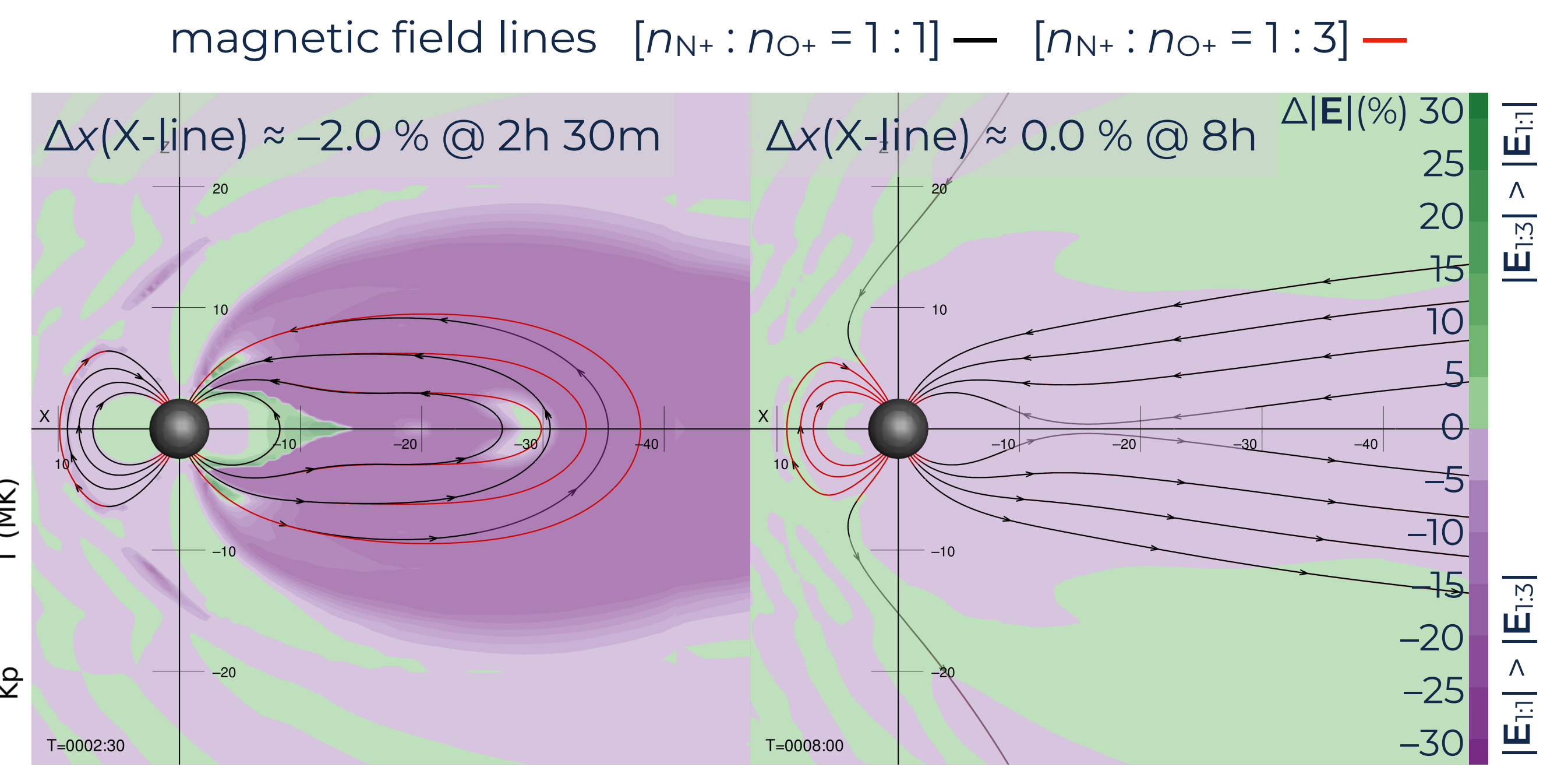
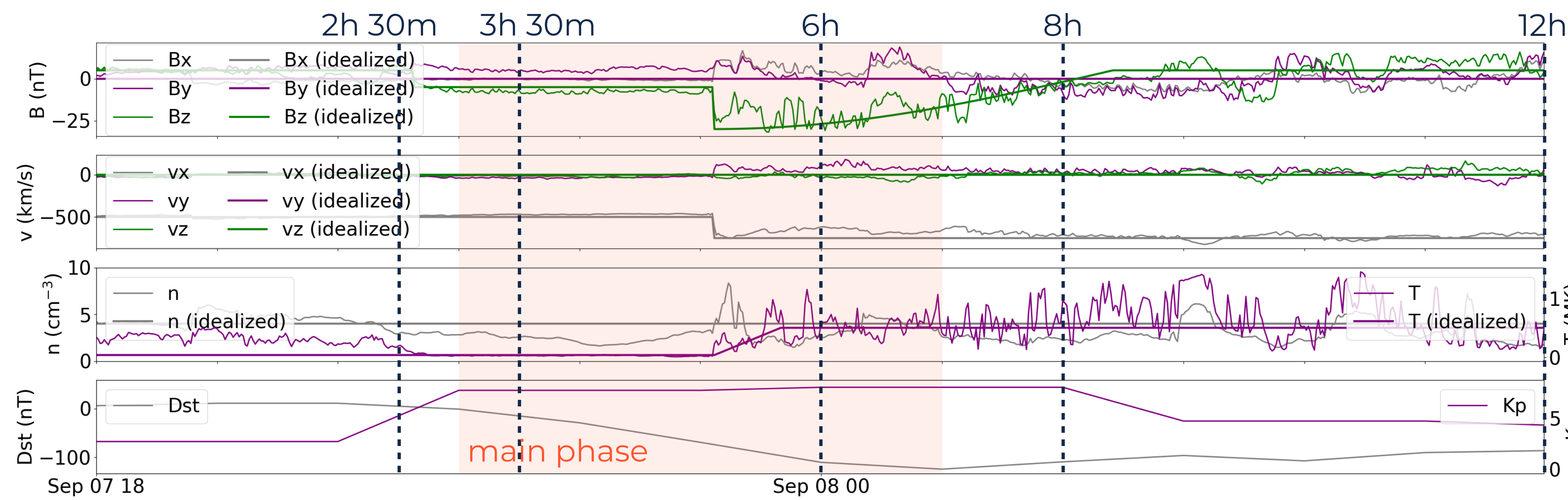
METHODOLOGY

- Multifluid Magnetohydrodynamics (MHD): H^+ , N^+ , O^+
- Space Weather Modeling Framework (SWMF)

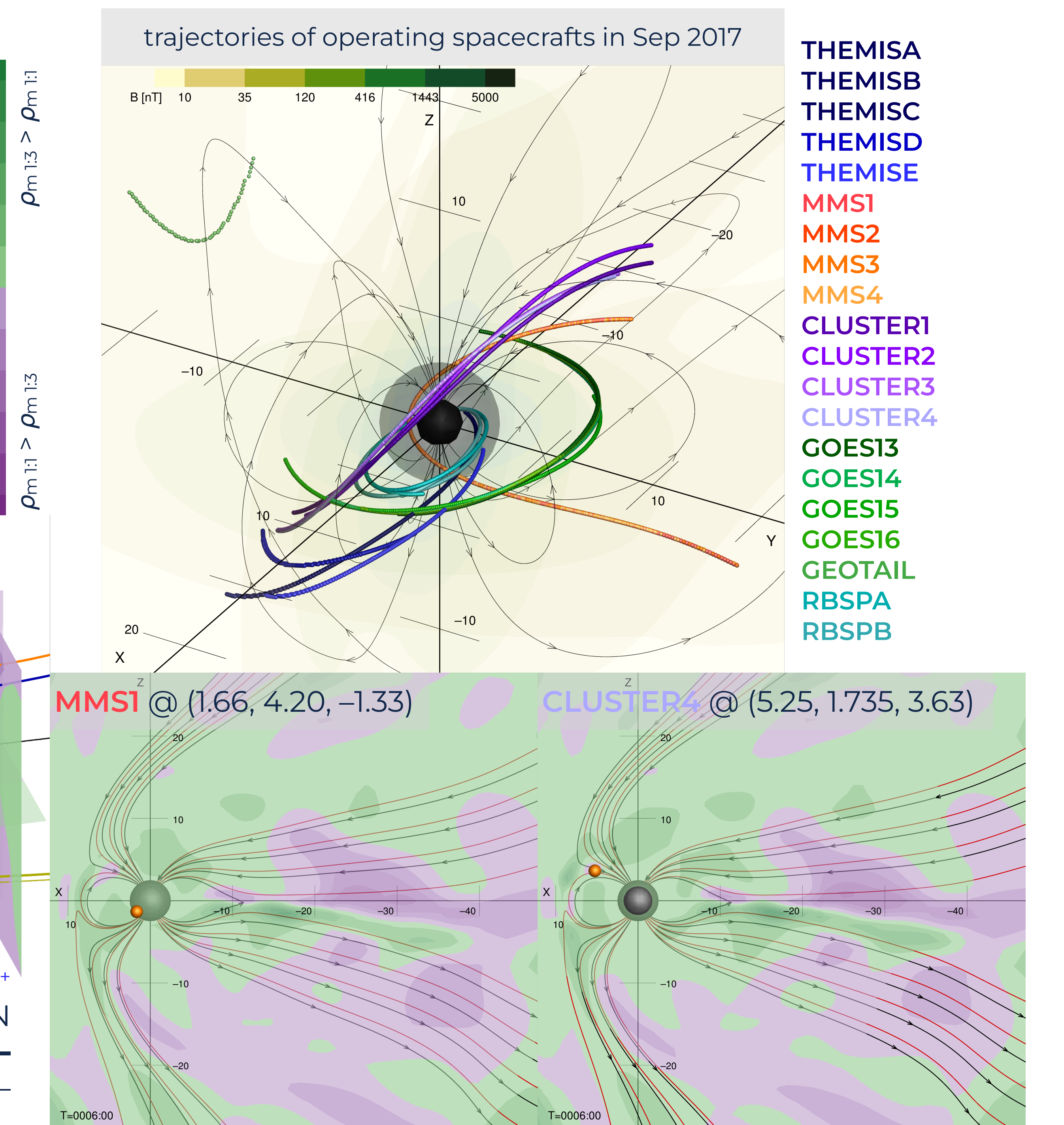
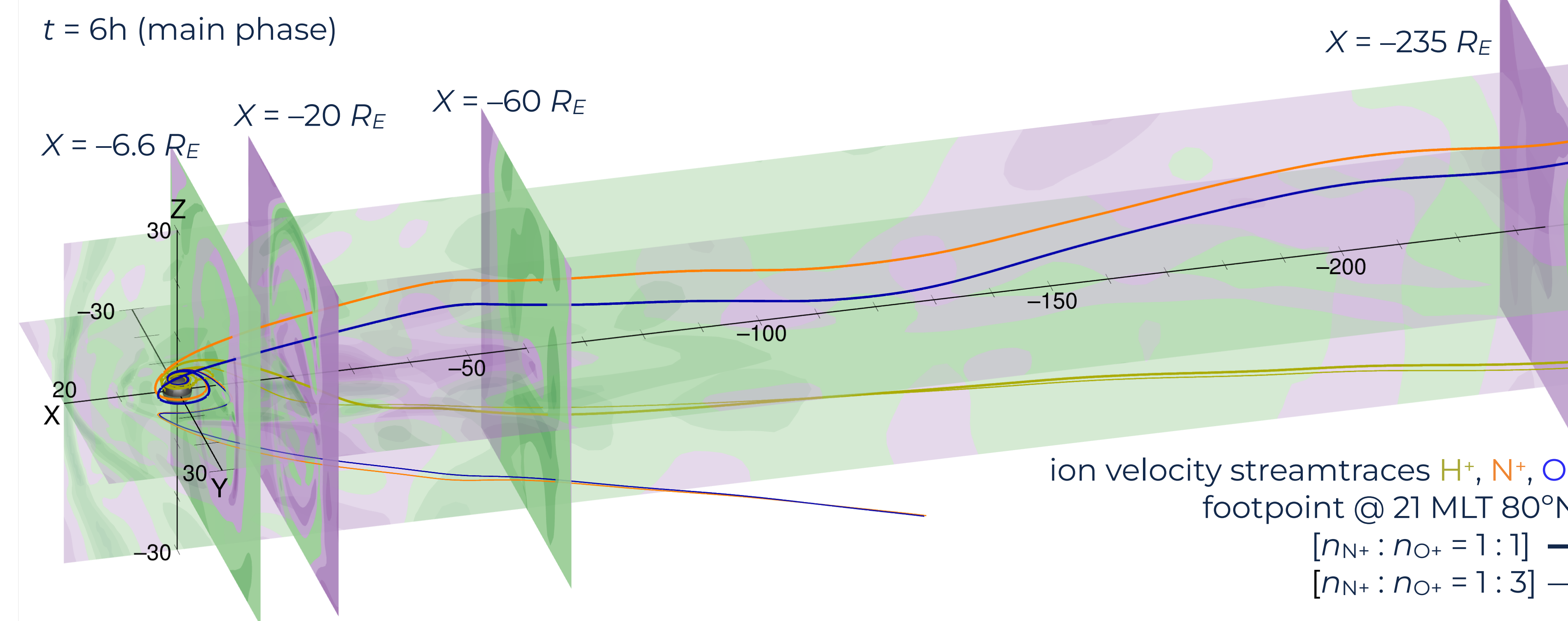
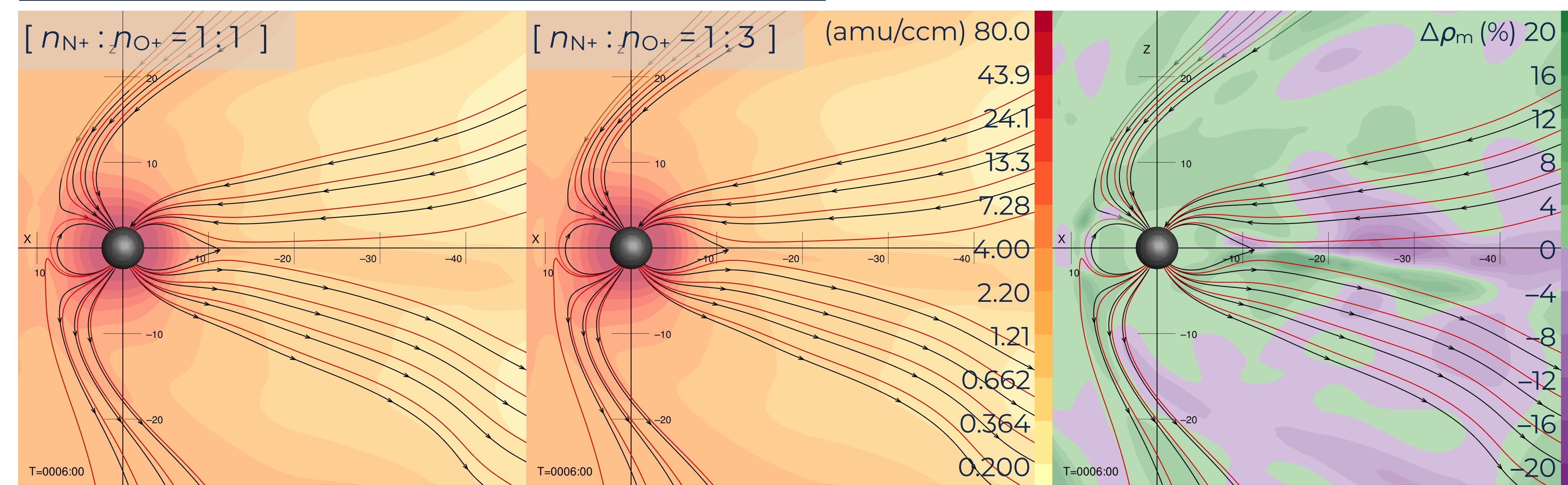


- 12-hour solar wind driver: DISCOVER measurements of the Sep 2017 storm (Sep 7 18:00 – Sep 8 6:00)
- Inner Boundary Ion Density Setup ($\sum_i \rho_i = 28 \text{ cm}^{-3}$):
 - $[n_{N^+} : n_{O^+} = 1:1]$ 80% H^+ , 10% N^+ , 10% O^+
 - vs. $[n_{N^+} : n_{O^+} = 1:3]$ 80% H^+ , 5% N^+ , 15% O^+ @ 2.5 R_E
- Domain: 292 $R_E \times 256 R_E \times 256 R_E$ (~2.6M cells)

IDEALIZED CONDITION: E & B FIELDS



STORM CONDITION: MASS DENSITY



CONCLUSION

- Even under idealized conditions, ion composition alters the electric and magnetic fields in the magnetosphere.
- Under the driving forces of observed solar wind data, the total mass density for case $[n_{N^+} : n_{O^+} = 1:1]$ is lower outside the magnetosphere and in the far tail, while the total mass density for case $[n_{N^+} : n_{O^+} = 1:3]$ is higher in the lobe regions in the magnetosphere.
- Heavy ion streamtraces differ greatly between the two cases.
- The magnetic field difference between the two cases is more prominent in observed conditions.

ACKNOWLEDGMENTS

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